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# REVISITING THE ROLE OF OBSERVATION SESSION DURATION ON PRECISE POINT POSITIONING ACCURACY USING GIPSY/OASIS II SOFTWARE

Rediscutindo o papel da duração da sessão de observação na acurácia do Posicionamento por Ponto Preciso com uso do software GIPSY/Oasis II

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#### Abstract:

The accuracy of GPS precise point positioning (PPP) was previously modelled as a function of the observing session duration T. The NASA, JPL's software GIPSY OASIS II (GOA-II) along with the legacy products was used to process the GPS data. The original accuracy model is not applicable anymore because JPL started releasing its products using new modelling and analysis strategies as of August 2007, and the legacy products are no longer available. The developments mainly comprise the new orbit and clock determination strategy, second order ionosphere modelling, and single station ambiguity resolution. Previously, the PPP accuracy was studied using v 4.0 of the GOA-II. The accuracy model showed coarser results compared to that of the relative positioning. Here, we processed the data of the International GNSS Service (IGS) stations to refine the accuracy of GOA-II PPP from version 6.3. Considering the above changes we refined the accuracy of PPP. First we modified the previous model used for the accuracy assessment. Then we tested out this model using straightforward polynomial and logarithmic models. The tests indicate the previous formulation still satisfactorily models the accuracy using refined coefficient values  $S_n = 7.8 \text{ mm}\sqrt{h}/\sqrt{T}$ ,  $S_e = 6.8 \text{ mm}\sqrt{h}/\sqrt{T}$ ,  $S_v = 29.9 \text{ mm}\sqrt{h}/\sqrt{T}$  for  $T \ge 2$  h.

**Keywords:** GPS accuracy, Precise Point Positioning, GIPSY OASIS II, observing session duration.

#### Resumo:

A acurácia do posicionamento por ponto preciso (PPP) do receptor GPS foi previamente modelada como uma função de observação com duração de tempo T. Para processar os dados

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GPS foram empregados os softwares GIPSY OASIS II (GOA-II) da JPL/NASA. Como a JPL passou a usar uma modelagem e novas estratégias de análises, o modelo convencional não é mais aplicável. A nova modelagem aborda, principalmente, novas órbitas e estratégias de determinação dos relógios dos satélites, modelagem de segunda ordem da ionosfera, e uma resolução única de ambiguidade. A acurácia do PPP foi previamente estudada usando a versão 4.0 do GOA-II. A acurácia do modelo mostrou resultados grosseiros comparados ao posicionamento relativo. Neste trabalho, foram processados os dados do serviço da estação internacional GNSS (IGS) para refinar a acurácia do GOA-II no modo PPP. Primeiramente, foram modificados os modelos prévios usados para avaliação da acurácia. Em seguida, os modelos foram testados empregando funções polinomiais simples e modelos logarítmicos. Os resultados mostraram que a formulação prévia ainda modela satisfatoriamente a acurácia usando os seguintes valores de coeficientes refinados:  $S_n = 7.8 \text{ mm}\sqrt{h}/\sqrt{T}$ ,  $S_e = 6.8 \text{ mm}\sqrt{h}/\sqrt{T}$ ,  $S_v = 29.9 \text{ mm}\sqrt{h}/\sqrt{T}$  para T > 2 h.

**Palavras-chave:** acuracidade GPS, Posicionamento por Ponto Preciso, GIPSY OASIS II, duração da sessão de observação.

### **1. Introduction**

Researchers and practitioners would like to know about the accuracy of GPS. This is needed for the accurate determination of deformation rates (Akarsu et al., 2015) and hence for survey planning. The use of PPP in real time GPS applications such as early prediction of earthquakes recently gained importance (Write et al., 2012). From this point of view, it is useful to revise the achievable accuracy of PPP from well-established GPS software, GIPSY/OASIS II (GOA-II) (Webb and Zumberge, 1997).

This area is of interest to researches for more than two decades now. Various studies were performed in regard to the assessment of the GPS accuracy such as by Davis et al. (1989), Dong and Bock. (1989), Larson and Agnew (1991), Heflin et al. (1992), Feigl et al. (1993), Rius et al. (1995), Eckl et al. (2001), Sanli and Engin (2009), Sanli and Tekic (2010), Bertiger et al. (2010), and Firuzabadi and King (2012). While some of these groups merely summarized the accuracy from their particular GPS experiments (i.e., usually from crustal deformation studies), the others took the accuracy of GPS as the main focus for their study and developed prediction formulas. The accuracy was found to be dependent on various factors such as baseline length, observing session duration, and network geometry. The software used in these studies also played a role. The results were assessed both for relative positioning (RP) and precise point positioning (PPP).

One of the first accuracy formulations, among the RP studies, was implemented by a National Geodetic Survey (NGS) team using the CORS stations of the country (Eckl et al., 2001). Since the stations were separated in 25-300 km baseline lengths, the team tested relative positioning accuracy depending on baseline lengths as well as observing session durations. At these scales, using their in house software PAGES, the dependency on baseline lengths was found to be negligible. Hence the accuracy of relative positioning was entirely dependent on observing session durations. The prediction formulae produced by Eckl et al. (2001) was given as below:

$$S_n(T) = \frac{9.5mm\sqrt{h}}{\sqrt{T}}$$

$$S_e(T) = \frac{9.9mm\sqrt{h}}{\sqrt{T}}$$
(1)

$$S_{\nu}(T) = \frac{36.5mm\sqrt{h}}{\sqrt{T}}$$

where  $S_n$ ,  $S_e$ , and  $S_v$  are the accuracies of position components north, east and vertical in mm. *T* is the observing session duration.

Using the formulas of Eckl et al. (2001), one was able to plan their survey prior to the field work. However their accuracy model was only appropriate for predicting the accuracy for 4 through 24 h data sets. Soler et al. (2006) developed an additional model that predicts the accuracy of 1-3 h sessions. Then the resulting formulation was:

$$s(cm) = \sqrt{\left(\frac{RMS_o}{1.5}\right)^2 \times \left(\frac{k}{\sqrt{T}}\right)^2 + \left(\frac{p}{1.7}\right)^2} \tag{2}$$

where s denotes the accuracy for short session durations;  $RMS_0$  (cm) is the overall RMS for the double differenced ionosphere-free carrier phase observables for the three single baseline solutions; k take on the values  $1 \text{ cm}\sqrt{h}$ ,  $1 \text{ cm}\sqrt{h}$ , and  $3.7 \text{ cm}\sqrt{h}$  for the north, east, and vertical components respectively; and p represents peak-to-peak errors along the north, east and vertical components (see Soler et al., 2006, for details).

Firuzabadi and King (2012) later on tested the effect of network geometry, i.e. distribution and number of network points, as well as the effect of observing session duration on the positioning accuracy. They found that the current IGS accuracy can only be achievable with 6 h of GPS data and 4 or more well distributed reference stations.

PPP refers to the mm-cm positioning of a GPS station using a single receiver, and was developed by NASA's Jet Propulsion Laboratory in mid 90s and then became a robust alternative to the traditional relative positioning technique (Zumberge et al., 1997). The user needs to use NASA's precise orbits and clocks as well as earth orientation parameters during the processing. Hence the positioning philosophy is actually based on a network approach in which a network of GPS stations is used to generate the so called products. The idea is to estimate precise satellite orbits from a core network of GPS stations, and not including a user's station to this estimation would hardly affect the quality of the orbits used in data processing. Obviously the PPP philosophy differs from the philosophy of Standard Point Positioning (SPP) in which only code measurements and the broadcast ephemeris are used to produce positioning information with a worst-case accuracy of about 3.5 m for horizontal and 4.5 m for vertical components (with 95% confidence as noted by www.gps.gov). PPP uses a combination of undifferenced code and carrier phase observations. Receiver clocks and tropospheric delays are estimated during the data processing. Then the solution is refined using sophisticated numerical algorithms and filtering techniques. However, the technique suffers from short observing session durations as indicated by Sanli and Engin (2009) and Sanli and Tekic (2010). If not modelled properly, the influences of global disturbances, which could easily be eliminated by relative positioning over shorter distances, are observed on the solutions (Blewit, 1993). However, the fact that one only needs a single receiver rather than minimum two is the advantage of the technique over relative positioning. The technique is also computationally efficient in that station-wise processing enables the processing time growing linearly.

Witchayang and Segantine (1999) and Fard and Dare (2006) were among the initial attempts assessing the PPP GPS results from web based GIPSY/OASIS software. Sanli and Engin (2009) noted the fact that practical applications such as monitoring of the tectonic motion and earthquakes might necessitate working over longer baseline lengths considering the need for using GPS stations from far stable tectonic plates such as Eurasia. Thus, for this concern in mind they revisited the accuracy of GPS positioning using GOA-II software. However, they did not use the PPP module as described in Zumberge et al. (1997), besides they fixed the ambiguities between station pairs using the ambiguity resolution technique of Blewitt (1989). This procedure produced results similar to those of the relative positioning. Namely, the dependency to GPS baselines was in concern. The results indicated that the positioning accuracy for the baselines longer than 300 km was dependent on baseline lengths. They produced new accuracy formulas in which the accuracy was dependent both on the observation session and baseline length:

$$S_n(L,T) = [80.19 \times 10^{-6} mm^2 h/km^2 L^2/T]^{0.5}$$

$$S_e(L,T) = [218.21 \times 10^{-6} mm^2 h/km^2 L^2/T]^{0.5}$$

$$S_v(L,T) = [951.83 \times 10^{-6} mm^2 h/km^2 L^2/T]^{0.5}$$
(3)

where L represents the baseline length between GPS stations.

Sanli and Tekic (2010) made an attempt to model (i.e. to formulate) the accuracy of GOA-II PPP applying ordinary PPP processing, and they did not fix ambiguities between station pairs. The results were produced based on JPL's legacy products (i.e. precise orbits, clocks, earth orientation parameters, etc. which were produced prior to the release of the new and improved products based on enhanced modelling and analysis procedures). The version 4.0 of GOA-II software was used for the data processing. Unlike relative positioning; the RMS of the solutions from shorter sessions, especially the east and the vertical components, showed dependency on station latitude. Namely, as the unsigned station latitude became smaller, the RMS of the solutions became larger. However, they only produced a model in which the positioning accuracy is formulated as a function of observing session duration. This model does not produce realistic results today because legacy products are no longer available and the processing needs to be based on the new and improved products. The refinement of the available accuracy model is the first task of this study, and it will be discussed in Section 2, in detail.

As Sanli and Tekic (2010) were studying the accuracy of GOA-II PPP, a new improved orbit determination strategy was newly adopted by NASA, JPL. Furthermore, the models to eliminate the second order ionosphere, which might be effective on the positioning at equatorial latitudes, were not available among the GOA-II modules (Kedar et al., 2003). Then, a single station ambiguity solution by Bertiger et al. (2010) was introduced that mainly improved the accuracy of horizontal positioning. All these were structured starting from the GOA-II PPP version 5.0 onwards. These enhancements were not taken into account in the analysis of Sanli and Tekic (2010).

Here, considering all these developments we revisit the role of the observing session duration on GOA-II PPP and produce a new accuracy model. We adopted the same GPS network as used in Sanli and Tekic (2010). We also adopted a similar modelling strategy as done in Sanli and Tekic (2010) and modelled the accuracy of the PPP depending only on the observing session duration. Furthermore we tested out additional models in which the available complex accuracy Bol. Ciênc. Geod., sec. Artigos, Curitiba, v. 22, nº,3 p.405-419, jul-set, 2016.

formulation is replaced with straightforward polynomial and log functions. IGS stations used in the study are almost evenly scattered around the globe depending on the data availability at the time. Then, we carried out the experiment using 2008 GPS data as done in Sanli and Tekic (2010). This was necessary for a direct comparison because the ionospheric conditions in 2008 coincided with the last solar cycle minimum whereas the present day data collection period coincides with a solar cycle maximum. The latest version 6.3 of GOA-II software was used in the processing of the GPS data (Desai et al., 2014). This version embraces improvements in orbit modelling, ambiguity resolution and ionosphere modelling. To show the enhancements from the latest version of GOA-II PPP, the results produced in this study are compared with those of the PPP studies produced earlier.

### 2. Original Accuracy Modelling

First of all, as also stated in Eckl et al. (2001), the accuracy here is not assessed by comparing GPS results with the results obtained from other space techniques such as VLBI and SLR. We also do not assess the accuracy in comparison to other software. The accuracy pronounced here is obtained by comparing the solutions from short GPS observation sessions with those of 24 h sessions, assuming the best positioning accuracy and hence the corresponding result is obtained through using 24 h data. Hence, the average of 24 h solutions is taken as the truth for the computation of the RMS from solutions of the shorter observation sessions. Therefore here the term "accuracy" rather than "precision" is used in assessing PPP results statistically.

In this context, the accuracy of PPP from JPL legacy products using GOA-II 4.0 was previously formulated as in the following (Sanli and Tekic, 2010):

$$S_n(\varphi, T) = \left[\frac{a_n}{T} + \frac{b_n \varphi^2}{T} + c_n + d_n \varphi^2\right]^{0.5}$$
(4)

where  $S_n(\varphi, T)$  denotes the accuracy of the north position component in mm as a function of station latitude  $\varphi$  in degrees and observation session length *T* in hours. The coefficients  $a_n$ ,  $b_n$ ,  $c_n$ ,  $d_n$  are derived through a least squares analysis (Eckl et al., 2001). Equation 4 also applies for the position components east and vertical.

Sanli and Tekic (2010) noticed the effect of station latitude on positioning results, especially on short observation sessions (i.e. 1-4 h) however they only modelled PPP results from 6-24 h GPS. Therefore, only the coefficient a , i.e. the effect of observing session duration, was found to be significant for all three GPS coordinates north, east and vertical as a result of the least squares analysis. Hence the accuracy of GPS PPP was formulated as

$$S_n(T) = \frac{13.5 \text{ mm} \cdot \sqrt{h}}{\sqrt{T}}$$

$$S_e(T) = \frac{20.7 \text{ mm} \cdot \sqrt{h}}{\sqrt{T}}$$

$$S_v(T) = \frac{40.8 \text{ mm} \cdot \sqrt{h}}{\sqrt{T}}$$
(5)

We first tested the modelling of Sanli and Tekic (2010) in which the modelling was merely performed with respect to observation session duration leaving the latitudinal effect out of the scope of the study. In fact, Sanli and Tekic (2010) used a model similar to the one given in Eckl et al. (2001). The only difference was that in their formula the dependency on station latitudes rather than baseline lengths was considered. Note that since they were using an absolute positioning technique (PPP) their processing was affected by global disturbances such as ocean, atmospheric and ground water loading rather than local or regional ones as given in Eckl et al. (2001).

We used the same IGS stations and same GPS data as used in Sanli and Tekic (2010) as also mentioned before. Table 1 gives the schedule of these observations.

**Table 1:** Schedule of observations where X denotes that data from this day were included in the study.

| Site<br>ID | Site<br>Latitude<br>(°) | Day of the Year (2008) |   |   |   |    |    |    |    |    |    |    |
|------------|-------------------------|------------------------|---|---|---|----|----|----|----|----|----|----|
|            |                         | 1                      | 2 | 8 | 9 | 10 | 11 | 12 | 13 | 14 | 15 | 16 |
| BOGT       | 4.6                     | х                      | х | х | х | х  | х  | х  | х  | х  | х  |    |
| YELL       | 62.5                    | х                      | х | х | х | х  | х  | х  | Х  | х  | х  |    |
| HARB       | -25.0                   | х                      | х | х | х | х  | х  | х  | х  | х  | х  |    |
| GUAM       | 13.6                    | х                      | х | х | х | Х  | х  | х  | х  | Х  | х  |    |
| MCM4       | -77.8                   | х                      | х | х | х | Х  | х  | х  | х  | х  | х  |    |
| MALI       | -3.0                    | х                      | х | х | х | Х  | х  | Х  | Х  | Х  |    | х  |
| IRKT       | 52.2                    | х                      | х | х | х | Х  | х  | Х  | Х  | Х  | Х  |    |
| QUIN       | 40.0                    | х                      | х | х | х | х  | х  | х  | х  | х  | х  |    |
| RIGA       | 57.0                    | х                      | х | х | х | х  | х  | х  | х  | х  | х  |    |
| MORP       | 55.2                    | х                      | х | х | х | х  | х  | х  | х  | х  | х  |    |
| SFER.      | 36.5                    | х                      | х | х | х |    | х  | х  | х  | х  | х  | х  |

## **3. PPP Processing**

We used GPS observations of the IGS stations scattered around the world downloading the data from Scripps Orbit and Permanent Array Centre (SOPAC) archives (sopac.ucsd.edu). We used the same GPS stations and same GPS data as given in Sanli and Tekic (2010). They are shown in Table 1. This way we were able to assess the accuracy improvement using the new and improved JPL orbits and clocks as well as IGS GPS data. The distribution of the IGS stations is shown in Figure 1.



Figure 1: Distributions of IGS permanent GPS stations used in the study.

With the motivation from the earlier studies we evaluated the effect of observing session duration on positioning solutions (Eckl et al., 2001; Sanli and Engin, 2009; Sanli and Tekic, 2010). In order to handle that, we sub-segmented 24 h observations into 1, 2, 3, 4, 6, 8, and 12 h sessions. From each session we computed a solution and then formed RMS values from the solutions of individual sub-sessions. The contaminated solutions, i.e. positional outliers, were removed using the most reliable robust outlier detection method 'median' as described in Tut et al. (2013). Solutions were also obtained from 24 h sessions and the average of those solutions was taken as the truth for computing the RMS of the shorter sessions.

The analysis of Sanli and Tekic (2010) did not include improvements from the NASA's new orbit processing strategy. JPL stopped generating legacy orbits in August 2009. Sanli and Tekic (2010) analysis also did not account for the improvements due to new single station ambiguity resolution technique as developed by Bertiger et al. (2010). Furthermore, second order ionosphere effects were not taken into account since the related module based on the modelling of Kedar et al. (2003) was first released with the version 5.0.

Apart from the new developments listed above we performed our analysis with the details given below. GPS data were obtained in Receiver Independent Exchange (RINEX) format. The data were sampled with 7.5 degree elevation cut off angle and 30 second recording intervals. JPL final precise (flinnR) orbits and clocks were used in the analysis. GIPSY/OASIS software's "precise point positioning" (PPP) was developed by Zumberge et al. (1997). In this study, the latest version 6.3 of the software was used. GIPSY utilizes precise satellite orbit and clock data from a global network solution (i.e. JPL "final" precise orbit determination). The results were represented using the International Earth Rotation Service's reference system ITRS (Petit and Luzum, 2010), as realized through the reference frame ITRF2008 (Altamimi et al., 2011). Tropospheric Zenith Wet Delay was modelled as a random-walk parameter with a variance rate of 5 mm2 per hour, and wet delay gradient as a random walk with a variance rate of 0.5 mm2 per hour. The dry troposphere was modelled using GMF mapping function a priori zenith conditions (Boehm et al., 2006). The ionospheric delay was eliminated from L1 and L2 data combination using both pseudo-range and carrier phase observations. Second order ionospheric correction was applied using the model of Kedar et al. (2003). Satellite and receiver antenna phase centre

variation (APV) maps were automatically applied following the IGS standards (Haines et al., 2010). Ocean loading corrections were also applied (Bos and Scherneck, 2015).

### 4. The New Accuracy Modelling

We first tested the mathematical model developed by Sanli and Tekic (2010) given with Equation 4 in Section 2. RMS values obtained from 24 h and all other sub-sessions for the GPS position components north, east and vertical are plotted against the observing session duration and given in Figure 2. Here note that as the observing session duration grows, the RMS of the solution becomes smaller. We will discuss later the meaning of the black, yellow, and red curves which are also superimposed on to the RMS values in Figure 2. The RMS values shown in Figure 2 were modelled for all three GPS position components north, east and vertical using the general relation given in Equation 4.

The improvement in JPL and IGS products led us also involve 4 h solutions in the modelling. Hence, differing from that of Sanli and Tekic (2010), which performed their modelling over 6-24 h session lengths, we performed our modelling using RMS values from 4-24 h solutions.

The coefficients *a*, *b*, *c*, and *d* for the each component was estimated using least squares. For the each constant of each GPS coordinate appearing in Equation 4, test statistics were formed by dividing the estimate to its formal error. These test statistics were then compared with the critical values obtained from Student's t test for  $\alpha = 0.05$  significance level.

The coefficients, i.e. the unknown parameters, less than the critical value 2.00 were eliminated from least squares solution (see Eckl et al., 2001, for details). As a result of the evaluation, only the coefficient a, i.e. the one merely depending on the observing session duration, was found to be significant. The statistics in regard to the constant a is given in Table 2.



**Figure 2:** Prediction/model fits to the RMS values of this study from observing session durations of 1 through 24 h (1-3 h period is just extrapolated). The model fit for the polynomial is represented with solid red line, for the log function with a dashed yellow line, and for Sanli and Tekic (2010) model with a black line.

**Table 2:** Significant constants/estimates and their 1-sigma uncertainties from least squares analysis for position components north, east, and vertical.

| Coefficient | Estimate | Uncertainty | Ratio |
|-------------|----------|-------------|-------|
| (mm²·h)     |          |             |       |
| $a_n$       | 60.31    | 10.72       | 5.63  |
| ae          | 44.97    | 5.35        | 8.40  |
| $a_v$       | 894.01   | 103.50      | 8.64  |

From these parameters, the new accuracy model based on the formulation of Sanli and Tekic (2010) was reformulated as below:

$$S_n = \frac{7.8 \text{ mm} \cdot \sqrt{h}}{\sqrt{T}}$$

$$S_e = \frac{6.8 \text{ mm} \cdot \sqrt{h}}{\sqrt{T}}$$

$$S_v = \frac{29.9 \text{ mm} \cdot \sqrt{h}}{\sqrt{T}}$$
(6)

where  $S_n$ ,  $S_e$ ,  $S_v$  are the accuracy estimates and *T* is the observing session duration. Comparing Equation 6 with Equation 5 reveals the improvement in the accuracy of GPS PPP mainly induced from JPL's new orbit determination strategy, ambiguity resolution technique and  $2^{nd}$  order ionosphere modelling. Namely, the model coefficients in Equation.5 have improved about 42% for the north, 67% the east, and 27% for the vertical coordinates. The solid black curve with triangular markers superimposed through the RMS values in Figure 2 represents the model derived using Equation 6.

#### **5.** Discussion

#### 5.1 Polynomial and Logarithmic Fitting

In this section, the accuracy of PPP will also be modelled using polynomial and logarithmic functions. This is because these functions are more straightforward, or less complicated than that of Sanli and Tekic (2010). In addition, Sanli and Tekic (2010) model only comprised the accuracy for the session durations of 4-24 h. Sessions 1-3 h in Figure 2 were only extrapolated. Note also that the segment of the solid black line fitting to 1 through 4 h sessions does not represent well the spectrum of the RMS. One reason in fitting polynomial and log functions to the RMS to test out whether or not an improvement could be gained in the modelling of shorter sessions.

The second order polynomial and a log function fitted into the RMS values of the north, east, and vertical coordinates were:

$$P(T) = a_0 + a_1 T + a_2 T^2 \tag{7}$$

$$L(T) = a \times ln(T) + b \tag{8}$$

where P(T) and L(T) denote the RMS derived from the polynomial and logarithmic models respectively. Coefficients  $a_0$ ,  $a_1$ ,  $a_2$ , a, and b are estimated through least squares analysis, and Tindicates the observing session duration. We applied similar least squares analysis and statistical testing procedures for the polynomial and logarithmic models given in Equations 7 and 8.

### **5.2 Internal Model Assessment**

Table 3 compares the predicted model values with the mean RMS values obtained using the RMS values given in Figure 2. The comparisons are made for all three GPS coordinates north, east, and up. The model fit for the polynomial is represented with solid red line whereas the model fit for the log function is represented with a dashed yellow line in Figure 2.

Note that the polynomial and logarithmic models are only successful in predicting the accuracy of 4-24 h solutions. However, Sanli and Tekic (2010) model is able to predict 2-24 h solutions. On the other hand, all three models fail to predict the actual mean RMS of 1h solutions. The refined Sanli and Tekic (2010) model given with Equation 6 still appears to be the best model which satisfactorily predicts the mean RMS for 2-24 h sessions. Appendix, attachments, glossary, and additional material must be available preferably in PDF format in which it will be linked in the whole text. If this material is part of a file, it must be presented after the References and follow the same formatting instructions of the text.

| Table 3: | Comparison of the predicted accuracy from three different models with | the a | actual |
|----------|---|-------|--------|
|          | mean RMS  |       |        |

| Coordinate | Session | Mean RMS | Model      |             |              |  |  |
|------------|---------|----------|------------|-------------|--------------|--|--|
|            | (h)     | (mm)     | Polynomial | Logarithmic | $a/\sqrt{T}$ |  |  |
|            | 1       | 19.7     | 3.9        | 4.5         | 7.8          |  |  |
|            | 2       | 5.2      | 3.7        | 3.9         | 5.5          |  |  |
|            | 3       | 4.1      | 3.5        | 3.5         | 4.5          |  |  |
|            | 4       | 3.4      | 3.3        | 3.2         | 3.9          |  |  |
| North      | 6       | 2.7      | 2.9        | 2.8         | 3.2          |  |  |
|            | 8       | 2.5      | 2.6        | 2.6         | 2.7          |  |  |
|            | 12      | 2.1      | 2.0        | 2.2         | 2.2          |  |  |
|            | 24      | 1.7      | 1.7        | 1.6         | 1.6          |  |  |
|            | 1       | 45.8     | 3.6        | 4.0         | 6.7          |  |  |
|            | 2       | 5.2      | 3.4        | 3.5         | 4.7          |  |  |
|            | 3       | 3.8      | 3.2        | 3.1         | 3.9          |  |  |
| _          | 4       | 3.0      | 3.0        | 2.9         | 3.4          |  |  |
| East       | 6       | 2.5      | 2.6        | 2.6         | 2.7          |  |  |
|            | 8       | 2.4      | 2.3        | 2.4         | 2.4          |  |  |
|            | 12      | 1.8      | 1.8        | 2.1         | 1.9          |  |  |
|            | 24      | 1.6      | 1.7        | 1.5         | 1.4          |  |  |
|            | 1       | 62.1     | 17.3       | 20.0        | 29.9         |  |  |
|            | 2       | 23.7     | 16.0       | 16.7        | 21.1         |  |  |
|            | 3       | 16.9     | 14.8       | 14.7        | 17.3         |  |  |
|            | 4       | 14.1     | 13.7       | 13.4        | 15.0         |  |  |
| Vertical   | 6       | 11.3     | 11.6       | 11.4        | 12.2         |  |  |
|            | 8       | 9.3      | 9.9        | 10.1        | 10.6         |  |  |
|            | 12      | 7.6      | 7.1        | 8.1         | 8.6          |  |  |
|            | 24      | 5.4      | 5.5        | 4.8         | 6.1          |  |  |

#### **5.3 External Model Assessment**

In this section we compare our 24 h predictions which are derived from Equation 6 with the evidence from previous studies. In addition, we also append the empirical evidence produced in this study and the current IGS positioning accuracies obtained from https://igscb.jpl.nasa.gov/ (Table 4).

**Table 4:** Comparison of the results from this study with those of the published work previously.Predictions of 24 h are compared.

| Accuracy                            | North<br>(mm) | East<br>(mm) | Vertical<br>(mm) |
|-------------------------------------|---------------|--------------|------------------|
| Refined accuracy model (Equation 6) | 1.6           | 1.4          | 6.1              |
| Average RMS                         | 1.7           | 1.6          | 5.4              |
| Bertiger et al. 2010                | 2.1           | 1.9          | 6.0              |
| Sanli and Tekic 2010                | 2.8           | 4.2          | 8.3              |
| IGS Continuous                      | 3.0           | 3.0          | 6.0              |

It is clearly seen from the table that Sanli and Tekic (2010) prediction is the poorest of all in representing the empirical evidence produced in this study. Note that the prediction of the east component from Sanli and Tekic (2010) became much better with the single receiver ambiguity solution developed by Bertiger et al. (2010). The refined second order ionosphere model and the required backward reprocessing by the JPL, including the corrections also for GPS satellites, came along with the software release v 6.3. The above mentioned enhancements in the software were lacking in the analysis of Sanli and Tekic (2010). Also note that 3-4 mm horizontal and 12 mm vertical GIPSY positioning accuracies using the model developed here and using as short as 4 h sessions do not yet satisfy the IGS accuracy given in Table 4.

Furthermore, the IGS vertical positioning accuracy is not even satisfied with 12 h of data. Akarsu et al. (2015) showed that the positioning using short observation sessions also affect the velocities estimated. Users of GPS campaign measurements employing 8-12 h of observation sessions need to take this into consideration.

### 6. Conclusion

The available accuracy model of GPS precise point positioning (PPP) was not reflecting the reality because the modelling was based on NASA, JPL's legacy products, and newly developed algorithms for data processing, such as single station ambiguity resolution and second order ionosphere modelling were not included. The user who wishes to use the previous accuracy model developed by Sanli and Tekic (2010) will not be able to obtain realistic results mainly because JPL's former legacy products are not distributed anymore. We refined the available accuracy model using the new and enhanced JPL and IGS products through GIPSY/OASIS II v 6.3 results (Desai et al. 2014). We also performed the modelling using straightforward

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polynomial and logarithmic functions. Comparisons from all three models revealed that with the refined model of Sanli and Tekic (2010), the accuracy can satisfactorily be predicted for session durations of 2 through 24 h. The accuracy estimated from the new model is comparable or shows improvement to those of the previous work. We assume the new model will be of interest to researchers who wish to plan their GPS campaigns prior to field works. As future work, we plan to include a non-linear coefficient c in our model (i.e.  $L(T) = a + \ln(T \times c) + b$ ) so that it might help improve the fit for 1 h sessions.

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